

# Trading linearity for ellipticity: a nonsmooth approach to Einstein's theory of gravity and Lorentzian splitting theorems

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[www.math.toronto.edu/mccann/Talk3.pdf](http://www.math.toronto.edu/mccann/Talk3.pdf)

7 May 2025

# What is a splitting theorem? (dimension reduction...)

## Example (convex functions, not necessarily smooth)

If the graph of a **convex** function  $u : \mathbf{R}^n \rightarrow \mathbf{R}$  contains a full line, say  $u(t, 0, \dots, 0) = 0$  for all  $t \in \mathbf{R}$ , then  $u(x) = U(x_2, \dots, x_n)$  for all  $x \in \mathbf{R}^n$

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## Example (smooth Riemannian manifolds; Cheeger-Gromoll '71)

If a connected **complete Ricci nonnegative** Riemannian manifold  $(M^n, g_{ij})$  contains a length minimizing line, then  $M$  is a geometric product of  $(\mathbf{R}, dr^2)$  with a submanifold  $(\Sigma^{n-1}, h_{ij} = g_{ij}|_{\Sigma})$ : i.e.

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**Gigli '13+**: nonsmooth version for **infinitesimally Hilbertian**

metric-measure spaces  $(M, d, m)$  satisfying **curvature-dimension** condition  **$RCD(0, N)$**  à la Sturm '06, Lott & **Villani** '09, (... M. '94)

**This talk**: **Lorentzian** analogs relevant to Einstein's theory of gravity

# Schematic proof of Cheeger-Gromoll splitting theorem:

Let  $\gamma : \mathbf{R} \rightarrow M^n$  be the doubly-infinite minimizing geodesic line.

Busemann '32:  $b_r(x) := d(x, \gamma(r)) - d(\gamma(0), \gamma(r))$  and  $\pm b^\pm := \lim_{r \rightarrow \pm\infty} b_r$

- note  $b_r$  is 1-Lipschitz and  $|\nabla b_r| = 1 = |\nabla b^\pm|$  a.e.; for  $r > 0$ ,
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- $\text{Ric} \geq 0$  gives **Hess**  $b = 0$  for  $b := b^\pm$
- hence  $\nabla b$  is a 'Killing' vector field (its **flow gives a local isometry**)
- $\Sigma := \{x \in M^n \mid b(x) = 0\}$  is totally geodesic (its **normal**  $\nabla b$  is **parallel**)
- along  $\Sigma$ , metric **splits** into **tangent**  $g_{ij} dy^i dy^j$  and **normal** components  $dr^2$
- $(r, y) \in \mathbf{R} \times \Sigma \mapsto \exp_y r \nabla b(y)$  is **surjective** hence a **global isometry**

# General relativity: Einstein's gravity and field equation

- Einstein's gravity is formulated on smooth Lorentzian manifolds, but often predicts such manifolds are geodesically incomplete (*nonsmooth*)

Gravity not a force, but rather a manifestation of curvature of spacetime

“Spacetime tells matter how to move” (along timelike/null geodesics... )

Field equation “Matter tells spacetime how to bend”

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$$\text{Ric}_{ij} - \frac{1}{2}Rg_{ij} = 8\pi T_{ij} \quad (\text{replaces } \Delta\phi = \rho \text{ and } F = -\nabla\phi)$$

- understanding Ricci and Einstein (à la Kip Thorne)  $i, j \in \{0, 1, 2, 3\}$

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- understanding Ricci and Einstein (à la Kip Thorne)  $i, j \in \{0, 1, 2, 3\}$
- just integrate this local conservation law for  $T_{ij}(x)$  to find  $g_{ij} \dots$

What if matter distribution is unknown?

Can look at initial value problem (nonlinear **wave** equation); linearization produces gravity waves; **no smoothing; singularities propagate**...

# Elliptic vs hyperbolic geometry

**ELLIPTIC:**  $\mathbf{R}^n$  equipped with Euclidean norm  $|v|_E := (\sum v_i^2)^{1/2}$

- $|v + w|_E \leq |v|_E + |w|_E$

**HYPERBOLIC:** Minkowski space  $\mathbf{R}^n$  equipped with the *hyperbolic 'norm'*

$$|v|_F := \begin{cases} (v_1^2 - \sum_{i \geq 2} v_i^2)^{1/2} & v \in F := \left\{ v \in \mathbf{R}^n \mid v_1 \geq (\sum_{i \geq 2} v_i^2)^{1/2} \right\} \\ -\infty & \text{else} \end{cases}$$

- $|v + w|_F \geq |v|_F + |w|_F$ , but terribly asymmetric

the *future*  $F \subset \mathbf{R}^n$  is a convex cone;  $v \in F$  called *causal* or *future-directed*

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- $v$  is *timelike* if  $v \in F \setminus \partial F$

- $v$  is *lightlike (or null)* if  $v \in \partial F \setminus \{0\}$

(•  $v$  is *spacelike* iff  $\pm v \notin F$  and *past-directed* if  $-v \in F$ )

- smooth *curves* are called *timelike (etc.)* if all tangents are timelike (etc.)



# A crash course in differential geometry: action principles

Manifold  $M^n$  with symmetric nondegenerate  $C^m$ -smooth tensor  $g_{ij} = g_{ji}$

**RIEMANNIAN:**  $(g_{ij}) > 0$  defines Euclidean norm on each tangent space

- its geometry is also encoded in the (symmetric) **distance** function

$$d(x, y) := \inf_{\sigma(0)=x, \sigma(1)=y} \left( \int_0^1 |\dot{\sigma}_t|_{E_g}^q dt \right)^{1/q} \quad q > 1$$

**LORENTZIAN:**  $g \sim (+1, -1, \dots, -1)$  defines hyperbolic norm on  $T_x M$

- its **asymmetric** geometry is also encoded in the **time-separation** function

$$\ell(x, y) := \sup_{\sigma(0)=x, \sigma(1)=y} \left( \int_0^1 |\dot{\sigma}_t|_{F_g}^q dt \right)^{1/q} \quad 0 \neq q < 1$$

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- $(-\infty)^q := -\infty$  so  $\ell(x, y) = -\infty$  unless a causal curve links  $x$  to  $y$
- extremizers are independent of  $q$ ; they are called **geodesics**
- $\ell(x, z) \geq \ell(x, y) + \ell(y, z)$  (analog of the **triangle inequality**  $d$  satisfies)

# The Riemann curvature tensor

Given (timelike) *geodesics*  $(\sigma_s)_{s \in [0,1]}$  and  $(\tau_t)_{t \in [0,1]}$  with  $\sigma_0 = \tau_0$  and  $\dot{\tau}_0, \dot{\sigma}_0, \dot{\tau}_0 - \dot{\sigma}_0 \in F \setminus \partial F$ , if  $g_{ij} \in C^3$

$$\ell(\sigma_s, \tau_t)^2 = |t\dot{\tau}_0 - s\dot{\sigma}_0|_{F_g}^2 - \frac{\text{Sec}}{6} s^2 t^2 + O((|s| + |t|)^5)$$

where sectional curvature  $\text{Sec} = R(\dot{\sigma}_0, \dot{\tau}_0, \dot{\sigma}_0, \dot{\tau}_0)$  is quadratic in  $\dot{\sigma}_0 \wedge \dot{\tau}_0$  and measures the leading order correction to Pythagoras

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- polarization of this quadratic form gives the *Riemann* tensor  $R(\cdot, \cdot, \cdot, \cdot)$
- its trace  $\text{Ric}_{ik} = g^{jl} R_{ijkl}$  yields the *Ricci* tensor;  $\text{Ric}(v, v)$  measures the correction to Pythagoras averaged over all triangles including side  $v$
- second trace  $R = g^{ik} \text{Ric}_{ik}$  yields the *scalar curvature*; in the elliptic case it gives leading order correction to the area of a sphere of radius  $r$  (and to the volume of a ball of radius  $r$ )
- $d\text{vol}_g(x) = \sqrt{|\det(g)|} d^n x$  in coordinates; (in the Riemannian case it coincides with the  $n$ -dimensional Hausdorff measure associated to  $d$ )

# Energy conditions and singularity theorems

- WEC** (weak energy condition):  $T(v, v) \geq 0$  for all **future**  $v \in F$  (**physical**)
- SEC** (strong energy condition):  $\text{Ric}(v, v) \geq 0$  for all **future**  $v \in F$  (**less** ")
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[Cosmological constant (dark matter):  $\geq (n-1)Kg(v, v)$ ]

THM: **Hawking '66** (big bang) singularities are generic:

**SEC** + mean curvature bound  $H_\Sigma \geq h > 0$  on a suitable hypersurface  $\Sigma$   
implies finite-time singularities along all **timelike** geodesics through  $\Sigma$

**Cavalletti-Mondino '20+**: also in  $TCD(0, N)$  metric-measure spacetimes

THM: **Penrose '65** (stellar collapse) singularities are generic

**NEC** + trapped **codimension-2 compact** surface  $S$  + suitable **noncompact hypersurface**  $\Sigma$  imply finite-time singularity along some **null** geodesic

**Graf '20** holds for  $g_{ij} \in C^1$ ;

**Open**: version for  $TCD(0, N)$  metric-measure spacetimes?

# Smooth Lorentzian splitting theorems

- '*spacetime*': a connected Lorentzian manifold  $(M^n, g_{ij})$  which admits a continuous choice of  $F_g$  (distinguishing future from past).
- '*strong energy condition*' *SEC*:  $g(v, v) > 0$  implies  $\text{Ric}(v, v) \geq 0$
- '*timelike geodesically complete*': all (unit speed) timelike geodesics admit doubly-infinite extensions (maximizing locally but not necessarily globally)

Theorem (conjectured by Yau '82; proved by Newman '90)

Let  $(M^n, g_{ij})$  be a SEC spacetime containing a maximizing timelike line. If  $M$  is (a) *timelike geodesically complete*, then  $M$  is a geometric product of  $\mathbf{R}$  with a  $(\text{Ric} \geq 0, \text{complete})$  Riemannian submanifold  $\Sigma^{n-1}$

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Galloway '84: proved assuming compact Cauchy surfaces exist

Beem, Ehrlich, Galloway, Markvorsen '85: proved under timelike sectional curvature nonnegativity assuming (b) global hyperbolicity;

Beran-Ohanyan-Rott-Solis '23: extension to g.h. Lorentzian length spaces

Eschenburg '88: proved under (a) + (b)

Galloway '89: proved under (b) without (a) using Bartnik '88



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- $(r, y) \in \mathbf{R} \times \Sigma \mapsto \exp_y r \nabla b(y)$  is surjective hence a global isometry

# Low regularity Eschenburg-Newman-Galloway theorem

- BGMOS quintet 24+ relax  $g_{ij} \in C^\infty(M^n)$  to  $g_{ij} \in C^1(M^n)$ , weights 25+  
Let  $\gamma : \mathbf{R} \rightarrow M^n$  be the isometrically embedded line and follow Busemann et al:  $b_r^+(x) := -\ell(x, \gamma(r)) + \ell(\gamma(0), \gamma(r))$  and  $b^\pm := \lim_{r \rightarrow \pm\infty} b_r^\pm$   
where  $b_r^-(x) := \ell(\gamma(r), x) - \ell(\gamma(r), \gamma(0))$
- reverse Lipschitz  $b_r^\pm(y) - b_r^\pm(x) \geq \ell(x, y)$  and  $|\nabla b_r^\pm|_F = 1$  a.e.;  $\forall r > 0$ ,
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## Theorem (BBCGMORS 24+ 'nonsmooth $p$ -d'Alembert comparison')

For  $p < 1$ , the operator  $\square_p u := -\nabla \cdot (|\nabla u|_F^{p-2} \nabla u)$  is **nonuniformly elliptic** and (SEC) implies  $\square_p b_r^+ \leq \frac{n-1}{\ell(\cdot, \gamma(r))}$  distributionally, i.e.  $\forall 0 \leq \phi \in C_c^1(M)$

# Low regularity Eschenburg-Newman-Galloway theorem

- **BGMOS quintet 24+** relax  $g_{ij} \in C^\infty(M^n)$  to  $g_{ij} \in C^1(M^n)$ , weights **25+**  
Let  $\gamma : \mathbf{R} \rightarrow M^n$  be the isometrically embedded line and follow Busemann  
et al:  $b_r^+(x) := -\ell(x, \gamma(r)) + \ell(\gamma(0), \gamma(r))$  and  $b^\pm := \lim_{r \rightarrow \pm\infty} b_r^\pm$   
where  $b_r^-(x) := \ell(\gamma(r), x) - \ell(\gamma(r), \gamma(0))$
- reverse Lipschitz  $b_r^\pm(y) - b_r^\pm(x) \geq \ell(x, y)$  and  $|\nabla b_r^\pm|_F = 1$  a.e.;  $\forall r > 0$ ,
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$$\int_M g \left( \nabla \phi, \frac{\nabla b_r^+}{|\nabla b_r^+|_F^{2-p}} \right) d\text{vol}_g \leq (n-1) \int_M \frac{\phi(\cdot) d\text{vol}_g(\cdot)}{\ell(\cdot, \gamma(r))}.$$

- for the distributional limit  $r \rightarrow \infty$ , need  $\nabla b_r^+ \rightarrow \nabla b^+$  strongly
- need **uniform ellipticity**; must bound  $\{\nabla b_r^+\}_{r \geq R}$  away from lightcone

# Convex $p$ -energy: trading linearity for ellipticity

Additional conditions may ensure  $\ell \neq +\infty$  and extremizers exist

- complete or *proper* (boundedly compact) in the Riemannian case
- (b) *globally hyperbolic* in the Lorentzian case (i.e. compact diamonds, future  $F$  varies continuously over  $M$ , no closed future-directed curves)

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satisfy  $DH = (DL)^{-1}$  if  $p^{-1} + q^{-1} = 1$  (here  $p < 0$  iff  $0 < q < 1$ )

- note  $L = H^*$  jumps from 0 to  $+\infty$  across future cone boundary  $\partial F$  (but  $H$  diverges continuously at the boundary of the dual cone  $F^*$ )
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**Beran Braun Calisti Gigli M. Ohanian Rott Sämann** (octet):

extremizers of  $p$ -Dirichlet energy  $u \mapsto \int_M H(du) d\text{vol}_g$  rel. to compactly supported perturbations satisfy a **new nonuniformly elliptic nonlinear** PDE

- trade linearity of d'Alembertian for ellipticity of  **$p$ -d'Alembertian!**

# Nondivergence expression of (nonuniform) ellipticity

$$\square_p b = \nabla_i \left( \frac{\partial H}{\partial w_i} \Big|_{db} \right) = H^{ij} \nabla_i \nabla_j b$$

$$H(w) = -\frac{1}{p} |w|_{F^*}^p$$

$$H^i := \frac{\partial H}{\partial w_i} = -|w|^{p-2} g^{ik} w_k$$

$$H^{ij} := \frac{\partial^2 H}{\partial w_i \partial w_j} = |w|^{p-2} \left[ (2-p) g^{ik} g^{jl} \frac{w_k w_l}{|w|^2} - g^{ij} \right]$$

$$\sim |w|^{p-2} \begin{bmatrix} 2-p-1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & \dots & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix} > 0 \quad \text{if } p < 1$$

choosing normal coordinates around  $\gamma(0)$  in which  $w = db$  is the time axis



Nomizu-Ozeki '61 give a complete Riemannian metric  $\tilde{g}$  on  $(M, g)$ .

Theorem (Eschenburg '88 ... Galloway-Horta '96)

Under (a) and/or (b),  $\gamma(0)$  admits a neighbourhood  $X$  and constants  $R, C$  such that if  $r \geq R$  then (i) a maximizing geodesic  $\sigma$  connects each  $x \in X$  to  $\gamma(r)$ ; (ii) each such geodesic satisfies  $\tilde{g}(\sigma'(0), \sigma'(0)) \leq Cg(\sigma'(0), \sigma'(0))$  hence  $\{b_r^+\}$  is timelike and uniformly *equiLipschitz* on  $X$ .

- intersecting ellipsoid and hyperboloid *uniformize ellipticity* on  $X$

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Lemma (BGMOS: *equi-semiconcavity* '24/-superdifferentiability '25+)

$g \in C^{1,1}$  yields  $\tilde{C}$  so all  $u \in \{b_r^+\}_{r \geq R}$  and  $(v, x) \in TX$  satisfy

$$\lim_{t \rightarrow 0} \frac{u(\exp_x^{\tilde{g}} tv) + u(\exp_x^{\tilde{g}} -tv) - 2u(x)}{\tilde{g}(v, v)} \leq \tilde{C}$$

(•  $p$ -d'Alembert comparison then follows from *smooth* calculations)

- gives  $db_r^+ \rightarrow db^+$  pointwise a.e., hence  $|db^\pm|_{F^*} = 1$  a.e. and
- $\pm b^\pm$  are distributionally  $p$ -superharmonic  $\square_p b^+ \leq 0 \leq \square_p b^-$

- now strong maximum principle improves  $b^+ \geq b^-$  to  $b^+ = b^- \in C^{1,1}(X)$  if  $g \in C^{1,1}$  (or  $b^\pm = b \in C^1(X)$  if  $g_{ij} \in C^1(M)$ )

Homogeneity  $2p - 2 < 0$  variant on Bochner Ohta '14, Mondino-Suhr '23:  
 $|du|_{F^*} = 1$  and  $\square_p u = 0$  imply (by elliptic regularity  $b \in W_{loc}^{2,2}(X)$  if  $g \in C^1(M)$ ), and

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- allow Bakry-Émery '85 weights, à la Case '10 and Woolgar-Wylie '18

## Theorem (BGMOS 25+): $C^1$ Eschenburg splitting with weight $V$ )

Let  $(M^n, g_{ij})$  with  $V, g_{ij} \in C^1$  be an (a) timelike geodesically complete, (b) globally hyperbolic spacetime and (c) timelike nonbranching, containing a maximizing timelike line. If  $N \in (n, \infty]$  and

$$\text{Ric}^{(N,V)} := \text{Ric} + \text{Hess } V - \frac{1}{N-n} dV \otimes dV \geq 0$$

timelike distributionally\* then  $M^n$  is ( $C^2$ -isometric to) a geometric product of  $(\mathbf{R}, dr^2)$  with a (complete,  $\text{Ric}^{(N-1,V)} \geq 0$  distributionally) Riemannian submanifold  $\Sigma^{n-1}$ . Also  $V(x(r, y))$  is independent of  $r \in \mathbf{R}$ .

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Proof: Set  $\nabla^{(g,V)} \cdot := (\nabla - dV) \cdot$  and  $\square_p^{(g,V)} u := -\nabla^{(g,V)} \cdot (|\nabla u|^{p-2} \nabla u)$ , and show Busemann functions satisfy  $\pm \square_p^{(g,V)} b^\pm \leq 0$  hence agree and are weighted  $p$ -harmonic. Use weighted Bochner-Ohta identity to deduce  $\text{Hess } b^\pm = 0$  and argue as before. □



THANK YOU VERY MUCH!