

Lecture 6

2026

Thm 2 (Bohlin) Assume $w \in \mathbb{C}$ moves according to floke $\ddot{w} = -w$. Let $z = w^2$, and new time τ is such that the law of sectorial area is satisfied. Then $z(\tau)$ satisfies Kepler: $\frac{d^2 z}{d\tau^2} = -c \frac{z}{|z|^3}$

Pf The law of areas: $\frac{|w|^2 d\varphi}{dt} = \text{const}$, $\frac{|z|^2 d\varphi}{d\tau} = \text{const}$
 Set $\frac{d\tau}{dt} = \frac{|z|^2}{|w|^2} \Rightarrow \frac{d}{d\tau} = \frac{|w|^2}{|z|^2} \frac{d}{dt} = \frac{1}{|w|^2} \frac{d}{dt}$ and

$$\begin{aligned} \frac{d^2 z}{d\tau^2} &= \frac{1}{w\bar{w}} \frac{d}{dt} \left(\frac{1}{w\bar{w}} \frac{dw^2}{dt} \right) = \frac{z}{w\bar{w}} \frac{d}{dt} \left(\frac{1}{w} \frac{dw}{dt} \right) \\ &= -\frac{z}{w\bar{w}} \left(\frac{1}{w^2} \frac{dw}{dt} \frac{d\bar{w}}{dt} + \frac{w}{\bar{w}} \right) = -2\bar{w}^{-1}\bar{w}^{-3} (|\dot{w}|^2 + |w|^2) \\ &\quad \nearrow \quad \quad \quad \nearrow \\ &= -4E \bar{w}^{-1}\bar{w}^{-3} \end{aligned}$$

used $\ddot{w} = -w$

Here $|\dot{w}|^2 + |w|^2 = 2E$ on the trajectory: conserv. of energy

$$\text{Since } \frac{z}{|z|^3} = \frac{w^2}{(w^2)^3} = \frac{w^2}{w^3 \bar{w}^3} = w^{-1} \bar{w}^{-3},$$

$$\Rightarrow \frac{d^2 z}{d\tau^2} = -c \frac{z}{|z|^3} \text{ for } c=4E \quad \blacksquare$$

Cor. Thm 1+2 \Rightarrow Kepler's bounded orbits are ellipses
w/ foci at 0 (i.e. for $E < 0$)

Thm 3 (Exer) Trajectories for force $\sim |w|^a$ are taken
by $z = w^\alpha$ for $\alpha = \frac{a+3}{2}$ into trajectories

$$w/\text{force} \sim |z|^A$$

(Find the relation of a and A)

Here $\overset{\text{Hooke}}{a=1}$, $\alpha = \frac{a+3}{2} = \frac{1+3}{2} = 2$, $z = w^2$

$\underset{\text{Kepler}}{A=-2}$

Similarity in physics

Observation: Let $\vec{r}(t)$ satisfies $m \frac{d^2 \vec{r}}{dt^2} = -\frac{\partial U}{\partial \vec{r}}$

take $t_1 = \alpha t$
 $m_1 = \alpha^2 m \mid \Rightarrow \vec{r}(t_1)$ satisfies $m_1 \frac{d^2 \vec{r}}{dt_1^2} = -\frac{\partial U}{\partial \vec{r}}$

i.e. if $m_1 = \frac{m}{9} \leadsto t_1 = \frac{t}{3}$ (3 times as fast)

Let $V(r)$ be a homogeneous f'n of degree ν

i.e.
$$V(\alpha r) = \alpha^\nu V(r) \quad \forall \alpha > 0$$

Exer
Prove

If γ is an orbit, then $\alpha\gamma$ is an orbit

What is the ratio of circulation times?

Derive $t_1 = t$ for Hooke's law $\nu = 2$, $U \sim r^2$

Find $t_1 = t(t)$ for $\nu = -1$, $U \sim \frac{1}{r}$

(obtain Kepler's III law)

Similarity

1. Kid's problem. Canada is 5,500 km wide

map is 55 cm

scale is $1 \text{ cm} = 100 \text{ km} = 100 \cdot \underset{\text{m}}{10^3} \cdot \underset{\text{cm}}{10^2} = 1 : 10^7$

Then ~ 40 mln population : $10^7 = 4$ persons should
stay on a map $1 : 10^7$. Catch? Areas $\sim (10^7)^2$


2. What is faster: to skin a bucket of large or small
(of the same weight) potatoes?

Fixed M , weight of 1 potat. $\sim \text{size}^3 = L^3$
 L -size skin area $\sim \text{size}^2 = L^2$

of potatoes in a bucket $n \sim \frac{1}{L^3} \Rightarrow$ total skin area $\sim n \cdot L^2 = \frac{L^2}{L^3} = \frac{1}{L}$

better - larger $L \Rightarrow$ less skin

3. A desert animal has to cover great distances w/o water. How does the max time the animal can run depend on its size L ?

Sol'n The storage of water \sim volume of body $\sim L^3$
evaporation \sim surface $= L^2 \Rightarrow$ the max time $\sim L$. 

Dimensional consideration $T = f(r, m, F)$

Physical units (following Maxwell) -

length, mass, time
 $L \quad M \quad t$

$$r \sim L, \quad m \sim M, \quad F = ma \sim M \cdot L / t^2$$

$$\Rightarrow \text{period } T \sim \left(\frac{m r}{F} \right)^{1/2}. \quad \text{Newton noted:}$$

$$\text{For } F = G \frac{m_1 m_2}{r^\alpha} \text{ with unknown } \alpha$$

$$\left(\frac{T_1}{T_2} \right)^2 = \left(\frac{m_1 r_1}{F_1} \right) / \left(\frac{m_2 r_2}{F_2} \right) = \left(\frac{\cancel{m_1} r_1}{\left(\frac{\cancel{m_1} M}{r_1^\alpha} \right)} \right) / \left(\frac{\cancel{m_2} r_2}{\left(\frac{\cancel{m_2} M}{r_2^\alpha} \right)} \right) = \frac{r_1^{\alpha+1}}{r_2^{\alpha+1}} = \left(\frac{r_1}{r_2} \right)^{\alpha+1}$$

Thus only $\alpha=2$ agrees w/ Kepler's III law!