# A fast algorithm for simulating scattering from a radially symmetric potential 

James Bremer

University of California, Davis

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The variable coefficient Helmholtz equation

$$
\Delta u(x)+k^{2}(1+q(x)) u(x)=0 \text { for all } x \in \mathbb{R}^{2}
$$

can be used to model the scattering of waves from inhomogeneous media.
In many applications:

- the wavenumber $k$ is real-valued
- the scattering potential $q$ is smooth, positive and has compact support contained in a disk $\Omega$ of radius $R$ centered at 0
- the solution $u$, which is known as the total field, is the sum of a known incident field $u_{i}$ satisfying the constant coefficient Helmholtz equation

$$
\Delta u_{i}(x)+k^{2} u_{i}(x)=0
$$

and an unknown scattered field $u_{s}$ which satisfies the Sommerfeld radiation condition

$$
\lim _{r \rightarrow \infty} \sup _{0 \leq t<2 \pi} \sqrt{r}\left|\frac{\partial u_{s}}{\partial r}\left(r e^{i t}\right)-i k u_{s}\left(r e^{i t}\right)\right|=0
$$

The goal is to obtain the scattered field $u_{s}$, which is the unique solution of the boundary value problem

$$
\left\{\begin{aligned}
\Delta u_{s}(x)+k^{2}(1+q(x)) u_{s}(x) & =-k^{2} q(x) u_{i}(x) \text { for all } x \in \mathbb{R}^{2} \\
\lim _{r \rightarrow \infty} \sup _{0 \leq t<2 \pi} \sqrt{r}\left|\frac{\partial u_{s}}{\partial r}\left(r e^{i t}\right)-i k u_{s}\left(r e^{i t}\right)\right| & =0
\end{aligned}\right.
$$

The scattered wave $u_{s}$ becomes increasingly oscillatory as $k$ grows, and it requires $N=\mathcal{O}\left(k^{2}\right)$ data points to represent it using standard methods (e.g., orthogonal polynomials, finite element bases, collocation).

As a consequence, any method which using such a scheme for representing $u_{s}$ has a running time which grows at least as fast as $N=\mathcal{O}\left(k^{2}\right)$.

In fact, most schemes have running times which grow superlinearly with $N$.

I will describe a method for calculating the scattered field in $\mathcal{O}(k \log (k))$ time.

The catch is that this method - at least in its present form - only applies to radially symmetric potentials $q$.

It operates by the method of separation of variables and exploits the fact that the one-dimensional variable coefficient Helmholtz equation

$$
y^{\prime \prime}(t)+k^{2} q(t) y(t)=0
$$

can be solved in $\mathcal{O}(\log (k))$ time.

## The one-dimensional variable coefficient Helmholtz equation

If $y(t)=\exp (r(t))$ solves

$$
\begin{equation*}
y^{\prime \prime}(t)+k^{2} q(t) y(t)=0, \tag{1}
\end{equation*}
$$

then $r$ satisfies the Riccati equation

$$
\begin{equation*}
r^{\prime \prime}(t)+\left(r^{\prime}(t)\right)^{2}+k^{2} q(t)=0 \tag{2}
\end{equation*}
$$

When $q$ is positive, the solutions of (1) are oscillatory, with the frequency of their oscillations increasing with $k$. Representing such functions using standard methods (e.g., expansions in orthogonal polynomials) requires $\mathcal{O}(k)$ data points.

Under mild conditions on $q$, (2) admits solutions which are nonoscillatory and can be represented via standard methods using a number of data points which is independent of $k$ (when $q$ is strictly positive or negative) and slowly growing with $k$ when $q$ has zeros.

## The one-dimensional variable coefficient Helmholtz equation

That solutions of the Riccati equation can be approximated by nonoscillatory functions is the basis of WKB approximation, and there are many examples from the theory of special functions of second order differential equations which have solutions whose logarithms are nonoscillatory.

For instance, the formula

$$
w(z):=\frac{1}{z} J_{\nu+\frac{1}{2}}^{2}(z)+\frac{1}{z} Y_{\nu+\frac{1}{2}}^{2}(z)=\frac{2}{\pi} \int_{0}^{\infty} \exp (-t z) P_{\nu}\left(1+\frac{t^{2}}{2}\right) d t
$$

shows that the function $w$ is completely monotone on $(0, \infty)$. It is related to a solution $r$ of the Riccati equation corresponding to the normal form of Bessel's differential equation via

$$
r^{\prime}(t)=i \frac{1}{w(z)}+\frac{1}{2} \frac{w^{\prime}(z)}{w(z)}
$$

## The one-dimensional variable coefficient Helmholtz equation

WKB approximations are not a numerically viable approach to computing nonoscillatory solutions of the Riccati equation because:

- they require the computation of high-order derivatives of the coefficient $q$
- the handling of turning points is messy

There is a viable algorithm, however, which runs in time independent of the wavenumber $k$ in the case in which $q$ is strictly positive and in $\mathcal{O}(\log (k))$ time in cases in which the equation has turning points:

B-, On the numerical solution of second order differential equations in the high-frequency regime. Applied and Computational Harmonic Analysis 44 (2018), 312-349.

It operates by solving the Riccati equation numerically. The only difficulty is computing the correct initial conditions.

## A numerical experiment

$$
q(t)=\frac{\nu^{2} \sin (2 t)^{2}}{0.1+(t-0.5)^{2}}+\mu^{2} \exp \left(-\frac{1}{t}\right)
$$

| $\nu$ | $\mu$ | Solve <br> time | Average <br> evaluation time |
| :---: | :---: | :---: | :---: |
| $10^{1}$ | $10^{0}$ | $1.79 \times 10^{-03}$ | $5.51 \times 10^{-07}$ |
| $10^{2}$ | $10^{0}$ | $3.10 \times 10^{-03}$ | $5.39 \times 10^{-07}$ |
| $10^{3}$ | $10^{0}$ | $2.90 \times 10^{-03}$ | $5.39 \times 10^{-07}$ |
| $10^{4}$ | $10^{0}$ | $3.06 \times 10^{-03}$ | $5.61 \times 10^{-07}$ |
| $10^{5}$ | $10^{1}$ | $3.80 \times 10^{-03}$ | $4.77 \times 10^{-07}$ |
| $10^{5}$ | $10^{2}$ | $4.62 \times 10^{-03}$ | $5.38 \times 10^{-07}$ |
| $10^{5}$ | $10^{3}$ | $3.64 \times 10^{-03}$ | $6.97 \times 10^{-07}$ |
| $10^{5}$ | $10^{4}$ | $3.81 \times 10^{-03}$ | $5.60 \times 10^{-07}$ |
| $10^{5}$ | $10^{5}$ | $3.63 \times 10^{-03}$ | $4.97 \times 10^{-07}$ |
| $10^{4}$ | $10^{5}$ | $2.72 \times 10^{-03}$ | $4.76 \times 10^{-07}$ |
| $10^{3}$ | $10^{5}$ | $3.57 \times 10^{-03}$ | $5.83 \times 10^{-07}$ |
| $10^{2}$ | $10^{5}$ | $3.15 \times 10^{-03}$ | $5.78 \times 10^{-07}$ |
| $10^{1}$ | $10^{5}$ | $2.32 \times 10^{-03}$ | $5.51 \times 10^{-07}$ |
| $10^{7}$ | $10^{7}$ | $4.33 \times 10^{-03}$ | $5.73 \times 10^{-07}$ |

## Representation of $u_{s}$ in the exterior of $\Omega$

Since $q$ is compactly supported in $\Omega$, the scattered wave $u_{s}$ is a solution of the boundary value problem

$$
\left\{\begin{align*}
\Delta u_{s}(x)+k^{2} u_{s}(x) & =0 \text { for all } x \in \Omega^{c}  \tag{3}\\
\lim _{r \rightarrow \infty} \sup _{0 \leq t \leq 2 \pi} \sqrt{r}\left|\frac{\partial u_{s}}{\partial r}(r, t)-i k u_{s}(r, t)\right| & =0
\end{align*}\right.
$$

We can separate variables in (3) to see that in the exterior of $\Omega, u_{s}$ can be represented as

$$
\begin{equation*}
u_{s}(r, t)=\sum_{n=-\infty}^{\infty} b_{n} H_{n}(k r) \exp (i n t) \tag{4}
\end{equation*}
$$

where $H_{n}$ is the Hankel function of the first kind of order $n$.

## Representation of $u$ in the interior of $\Omega$

The total field $u$ satisfies

$$
\Delta u(x)+k^{2}(1+q(r)) u(x)=0 \text { for all } x \in \Omega .
$$

Separating variables shows that $u$ can be represented in the interior of $\Omega$ via an expansion of the form

$$
\begin{equation*}
u(r, t)=\sum_{n=-\infty}^{\infty} a_{n} \psi_{|n|}(r) \exp (i n t) \tag{5}
\end{equation*}
$$

where, for each nonnegative integer $n, \psi_{n}$ is a solution of the second order differential equation

$$
r^{2} \psi_{n}^{\prime \prime}(r)+r \psi_{n}^{\prime}(r)+\left(k^{2}(1+q(r)) r^{2}-n^{2}\right) \psi_{n}(r)=0
$$

which is nonsingular at 0 . Note that $\psi_{n}$ is only determined up to a nonzero multiplicative constant, but the choice of this constant does not effect the form of (5), only the value of the coefficient $a_{n}$.

## Determination of the scattered field via separation of variables

From standard elliptic regularity results, we know that $u$ and its normal derivative are continuous across the boundary $\partial \Omega$. Moreover, $u=u_{s}+u_{i}$. It follows that

$$
\sum_{n=-\infty}^{\infty} a_{n} \psi_{|n|}(R) \exp (i n t)=\sum_{n=-\infty}^{\infty} b_{n} H_{n}(k R) \exp (i n t)+u_{i}(R, t)
$$

and

$$
\sum_{n=-\infty}^{\infty} a_{n} \psi_{|n|}^{\prime}(R) \exp (i n t)=\sum_{n=-\infty}^{\infty} b_{n} k H_{n}^{\prime}(k R) \exp (i n t)+\frac{\partial u_{i}}{\partial r}(R, t)
$$

## Determination of the scattered field via separation of variables

If we let

$$
u_{i}(R, t)=\sum_{n=-\infty}^{\infty} c_{n} \exp (i n t) \quad \text { and } \quad \frac{\partial u_{i}}{\partial r}(R, t)=\sum_{n=-\infty}^{\infty} d_{n} \exp (i n t)
$$

be the Fourier expansions of the restrictions of the incident wave $u_{i}$ and its derivative with respect to $r$ to the boundary $\partial \Omega$ of $\Omega$, then we can rewrite these equations as

$$
\sum_{n=-\infty}^{\infty} a_{n} \psi_{|n|}(R) \exp (i n t)=\sum_{n=-\infty}^{\infty} b_{n} H_{n}(R) \exp (i n t)+\sum_{n=-\infty}^{\infty} c_{n} \exp (i n t)
$$

and

$$
\sum_{n=-\infty}^{\infty} a_{n} \psi_{|n|}^{\prime}(R) \exp (i n t)=\sum_{n=-\infty}^{\infty} b_{n} k H_{n}^{\prime}(k R) \exp (i n t)+\sum_{n=-\infty}^{\infty} d_{n} \exp (i n t)
$$

## Determination of the scattered field via separation of variables

Owing to the orthogonality of the exponential functions, the equations

$$
\sum_{n=-\infty}^{\infty} a_{n} \psi_{|n|}(R) \exp (i n t)=\sum_{n=-\infty}^{\infty} b_{n} H_{n}(R) \exp (i n t)+\sum_{n=-\infty}^{\infty} c_{n} \exp (i n t)
$$

and

$$
\sum_{n=-\infty}^{\infty} a_{n} \psi_{|n|}^{\prime}(R) \exp (i n t)=\sum_{n=-\infty}^{\infty} b_{n} k H_{n}^{\prime}(k R) \exp (i n t)+\sum_{n=-\infty}^{\infty} d_{n} \exp (i n t)
$$

are satisfied if and only if for each integer $n$,

$$
\left\{\begin{aligned}
a_{n} \psi_{|n|}(R)-b_{n} H_{n}(R) & =c_{n} \\
a_{n} \psi_{|n|}^{\prime}(R)-b_{n} k H_{n}^{\prime}(k R) & =d_{n}
\end{aligned}\right.
$$

## Determination of the scattered field via separation of variables

We can easily solve this system:

$$
\begin{aligned}
& a_{n}=\frac{-k H_{n}^{\prime}(k R) c_{n}+H_{n}(k R) d_{n}}{H_{n}(k R) \psi_{|n|}^{\prime}(R)-k \psi_{|n|}(R) H_{n}^{\prime}(k R)} \\
& b_{n}=\frac{-\psi_{|n|}^{\prime}(R) c_{n}+\psi_{|n|}(R) d_{n}}{H_{n}(k R) \psi_{|n|}^{\prime}(R)-k \psi_{|n|}(R) H_{n}^{\prime}(k R)} .
\end{aligned}
$$

In the interior of $\Omega$ :

$$
u(r, t)=\sum_{n=-\infty}^{\infty} a_{n} \psi_{|n|}(r) \exp (i n t)
$$

In the exterior of $\Omega$ :

$$
u_{s}(r, t)=\sum_{n=-\infty}^{\infty} b_{n} H_{n}(k r) \exp (i n t) .
$$

## Numerical Algorithm

It is fairly straightforward to develop a numerical method based on the method of separation of variables.

We first observe that since the incident wave $u_{i}$ satisfies the constant coefficient Helmholtz equation at wavenumber $k$, we expect to be able to represent it and its normal derivative using $\mathcal{O}(k)$ Fourier modes. That is, in typical cases, we will have high accuracy approximations

$$
\begin{aligned}
u_{i}(t, R) & \approx \sum_{n=-m}^{m} c_{n} \exp (i n t) \\
\frac{\partial u_{i}}{\partial r}(t, R) & \approx \sum_{n=-m}^{m} d_{n} \exp (i n t)
\end{aligned}
$$

with $m=\mathcal{O}(k)$. Moreover, the coefficients in these expansions can be computed via the fast Fourier transform in $\mathcal{O}(m \log (m))$ operations.

## Numerical Algorithm

The number of terms in the expansion

$$
u_{i}(t, R) \approx \sum_{n=-m}^{m} c_{n} \exp (i n t)
$$

of the incident wave dictates the number of terms in the representations of total field $u$ and of the scattered field $u_{s}$.

For each $n=0, \ldots, m$, we must solve the perturbed Bessel equation

$$
r^{2} \psi_{n}^{\prime \prime}(r)+r \psi_{n}^{\prime}(r)+\left(k^{2}(1+q(r)) r^{2}-n^{2}\right) \psi_{n}(r)=0
$$

to compute $\psi_{|n|}$. The cost for each solve is $\mathcal{O}(\log (k))$, and $m=\mathcal{O}(k)$, so the total cost to compute the functions $\psi_{n}$ is $\mathcal{O}(k \log (k))$.

Computing the coefficients in the expansions of $u$ and $u_{s}$ requires $\mathcal{O}(k)$ operations.
So the total running time is $\mathcal{O}(k \log (k))$.

$$
q(r)=\exp \left(-5 r^{2}\right), \quad u_{i}(r, t)=\exp \left(i k r \cos \left(t-\frac{\pi}{4}\right)\right)
$$






| $k$ | $m$ | Maximum absolute <br> error | Precomp time <br> (in seconds) | Solve time <br> (in seconds) |
| :---: | ---: | :---: | :---: | :---: |
| $2^{8}$ | 1608 | $2.17 \times 10^{-12}$ | $4.60 \times 10^{-01}$ | $2.44 \times 10^{-02}$ |
| $2^{9}$ | 3216 | $4.72 \times 10^{-12}$ | $9.56 \times 10^{-01}$ | $1.06 \times 10^{-02}$ |
| $2^{10}$ | 6433 | $8.54 \times 10^{-12}$ | $1.96 \times 10^{+00}$ | $9.27 \times 10^{-02}$ |
| $2^{11}$ | 12867 | $1.85 \times 10^{-11}$ | $4.18 \times 10^{+00}$ | $2.22 \times 10^{-01}$ |
| $2^{12}$ | 25735 | $6.13 \times 10^{-11}$ | $8.99 \times 10^{+00}$ | $2.25 \times 10^{-02}$ |
| $2^{13}$ | 51471 | $2.05 \times 10^{-10}$ | $1.89 \times 10^{+01}$ | $1.52 \times 10^{-01}$ |
| $2^{14}$ | 102943 | $1.51 \times 10^{-09}$ | $4.08 \times 10^{+01}$ | $1.56 \times 10^{-01}$ |
| $2^{15}$ | 205887 | $3.85 \times 10^{-09}$ | $8.72 \times 10^{+01}$ | $2.32 \times 10^{-01}$ |
| $2^{16}$ | 411774 | $2.16 \times 10^{-08}$ | $1.93 \times 10^{+02}$ | $2.46 \times 10^{+00}$ |
| $2^{17}$ | 823549 | $1.01 \times 10^{-07}$ | $4.11 \times 10^{+02}$ | $1.36 \times 10^{+00}$ |




$$
q(r)=14 r^{2} \exp \left(-5 r^{2}\right), \quad u_{i}(z)=H_{0}(k|z-6 i|)
$$






| $k$ | $m$ | Maximum absolute <br> error | Precomp time <br> (in seconds) | Solve time <br> (in seconds) |
| :---: | ---: | :---: | :---: | :---: |
| $2^{8}$ | 1608 | $2.44 \times 10^{-12}$ | $4.61 \times 10^{-01}$ | $2.44 \times 10^{-02}$ |
| $2^{9}$ | 3216 | $5.85 \times 10^{-12}$ | $9.37 \times 10^{-01}$ | $1.03 \times 10^{-02}$ |
| $2^{10}$ | 6433 | $9.34 \times 10^{-12}$ | $1.98 \times 10^{+00}$ | $9.11 \times 10^{-02}$ |
| $2^{11}$ | 12867 | $1.98 \times 10^{-11}$ | $4.20 \times 10^{+00}$ | $2.27 \times 10^{-01}$ |
| $2^{12}$ | 25735 | $4.80 \times 10^{-11}$ | $9.00 \times 10^{+00}$ | $2.64 \times 10^{-02}$ |
| $2^{13}$ | 51471 | $4.35 \times 10^{-10}$ | $1.90 \times 10^{+01}$ | $1.49 \times 10^{-01}$ |
| $2^{14}$ | 102943 | $1.91 \times 10^{-09}$ | $4.09 \times 10^{+01}$ | $1.53 \times 10^{-01}$ |
| $2^{15}$ | 205887 | $6.93 \times 10^{-09}$ | $8.75 \times 10^{+01}$ | $2.34 \times 10^{-01}$ |
| $2^{16}$ | 411774 | $3.23 \times 10^{-08}$ | $1.93 \times 10^{+02}$ | $2.53 \times 10^{+00}$ |
| $2^{17}$ | 823549 | $1.08 \times 10^{-07}$ | $4.17 \times 10^{+02}$ | $1.13 \times 10^{+00}$ |




$$
q(r)=\chi_{(0,1)}(r)+2 \chi_{(2,3)}(r), \quad u_{i}(z)=\exp \left(i k r \cos \left(t-\frac{\pi}{4}\right)\right)
$$






| $k$ | $m$ | Maximum absolute <br> error | Time <br> (in seconds) |
| :---: | ---: | :---: | :---: |
| $2^{8}$ | 1608 | $\left(2.20 \times 10^{-12}\right)$ | $9.58 \times 10^{-01}$ |
| $2^{9}$ | 3216 | $\left(5.95 \times 10^{-12}\right)$ | $1.93 \times 10^{+00}$ |
| $2^{10}$ | 6433 | $\left(6.50 \times 10^{-12}\right)$ | $4.06 \times 10^{+00}$ |
| $2^{11}$ | 12867 | $\left(2.04 \times 10^{-11}\right)$ | $8.57 \times 10^{+00}$ |
| $2^{12}$ | 25735 | $\left(4.31 \times 10^{-11}\right)$ | $1.77 \times 10^{+01}$ |
| $2^{13}$ | 51471 | $\left(2.27 \times 10^{-10}\right)$ | $3.70 \times 10^{+01}$ |
| $2^{14}$ | 102943 | $\left(1.90 \times 10^{-10}\right)$ | $7.79 \times 10^{+01}$ |
| $2^{15}$ | 205887 | $\left(3.48 \times 10^{-10}\right)$ | $1.62 \times 10^{+02}$ |
| $2^{16}$ | 411774 | $\left(7.33 \times 10^{-10}\right)$ | $3.56 \times 10^{+02}$ |
| $2^{17}$ | 823549 | $\left(1.95 \times 10^{-09}\right)$ | $7.22 \times 10^{+02}$ |




## The general case?

In the two dimensional radially symmetric case, we represented the desired solution $u$ of

$$
\Delta u(x)+k^{2}(1+q(x)) u(x)=0
$$

in the form

$$
\sum_{n=-\infty}^{\infty} a_{n} \psi_{|n|}(r) \exp (i n t)
$$

with $\psi_{|n|}$ a solution of a second order differential equation whose logarithm is a nonoscillatory solution of the one-dimensional Riccati equation.

In fact, the logarithm $r$ of $\psi_{|n|}(r) \exp (i n t)$ is a solution of the two-dimensional Riccati equation

$$
\Delta r(x)+\nabla r(x) \cdot \nabla r(x)+k^{2}(1+q(x))=0
$$

## The general case?

In other words, we are representing the solution $u$ of

$$
\Delta u(x)+k^{2}(1+q(x)) u(x)=0
$$

as a sum of the form

$$
\sum_{n=-\infty}^{\infty} a_{n} \exp \left(\gamma_{n}(x)\right)
$$

where $\gamma_{n}$ is a nonoscillatory solution of the Riccati equation

$$
\Delta \gamma(x)+\nabla \gamma(x) \cdot \nabla \gamma(x)+k^{2}(1+q(x))=0
$$

such that $\exp \left(\gamma_{n}(x)\right)$ is a multiple of $\exp (i n \theta)$ on $\partial \Omega$.

